#### Features Detection following the SuperNova Early Warning System

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#### Motivations

- "Prediction" of Core-collapse supernova with its feature
- Things we can observe with photons
- Secondary burst
- Blackhole forming
- A possible evidence for quark star
- Provide information to the observer communities



# Background

- SuperNova Early Warning System
- Detection of supernova in our galaxy
- Neutrinos
- Core-collapse supernova
- Data from different detectors around Neutrino detector in SNOLAB the world
- Neutrinos reach a few hours before photons arrive



(picture from snolab.ca)

### Core-collapse supernova and neutrinos

- Powerful release a lot of energy
- 99% of the energy lost via neutrinos
- Rare event
- Last detection:

23 Feb, 1987





Remnant of SN1987A ESA/Hubble & NASA

Halliday, Resnick & Walker: Fundamentals of Physics

## Preparation

- Supernova
- A lot of information
- we missed a lot in 1987 lack of advance detectors/observers
- SNEWS detecting neutrinos
- The very first detection
- A few hours before photons reach
- Prepare all algorithms in advance
- Gather as much information as we can for the follow-ups (e.g. skymap)







## Possible features on supernova neutrinos

Secondary burst Computational simulations in different models - Oscillation Oscillation standing accretion shock instability  $10^{54}$ - Cut-off (erg/s) Turned into a blackhole Sudden rise J 10<sup>53</sup> (Supernova starts) - Secondary burst M32-S07 M32-S10M20-S05 M20-S07 Phase passing through Quark star? M20-S10  $10^{52}$ 10Cut off  $t_{\rm pb}$  (s) Fujibayashi et al., ApJ 919 (2021) 2, 80

## Simulation from another supernova model



Zha et al., ApJ 911 (2021) 2, 74

## Calculation

- Algorithm in python (following the SNEWS)
- As a plugin in the follow-up calculating system (Snewpdag)



- Information of supernova
- Distance, Direction (so that observers know where to point the telescope)
- And features!

## Detection of the features

- Neutrino detection is a rare event
- The incoming data is expected to be Poisson distributed
- Supernova is not the only source of neutrinos in the universe
- Background noise is expected

- It is likely that the detection may consist of errors and false alarms
- Target reduce the rate of these





# Algorithms (Secondary peak plugin)

- Simple version
- Detecting rising slope with threshold
- Input: Time series/histogram
- Output: True/False
- Problems:
- False alarms (low threshold)
- Missing features (high threshold)



# Algorithms (Secondary peak plugin)

- Utilize the property of Poisson distribution (ongoing)
- Input: Time series/histogram
- Output: Confidence level profile



- Preform calculation of the data from different detectors
- Compare them according to the relative time after the burst started
- Obtain the time point where second peak occur (with confidence level)

# Algorithms (cut-off plugin) By Teammate (Muhammad Hamza Kalim)

- Determine if the luminosity of neutrinos drops to background level
- Input: Time series/histogram
- Output: Confidence level value
- Compare the data from detectors



- Indicating that a blackhole is possibly formed
- All information is trapped no neutrinos will be further observed
- Failed supernova
- possible detection through the gravitational wave

NASA

## Testing the algorithm

- We need some data to test the accuracy
- Uniform generator + Gaussian generator
- Parameters (bg rate + total neutrinos events of a peak)
- Simple testing data source
- Able to be massively generated and tested
- Fairly accurate false/missing alarm rate



### Expectation

- Supernova occurs
- Get as much information as we can
- Alerts to the observers with description on the supernova to be seen



An artistic conception of a supernova explosion (NASA)

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