

Investigation of the Massive Infrared Dark Cloud in Early Star Formation

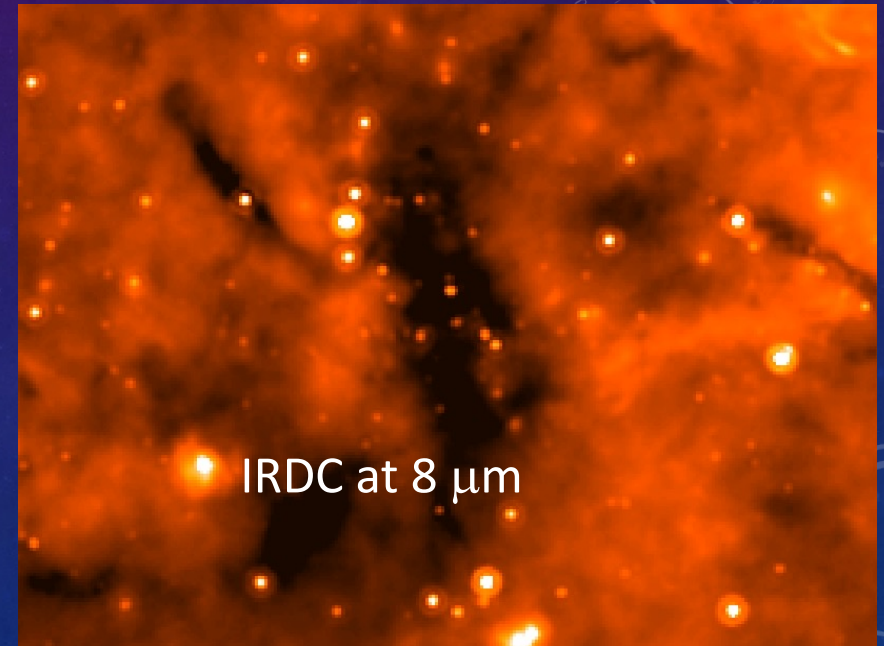
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Outline

- Introduction (IRDC, observation data)
- Method of Imaging
- Results
- Conclusion

Introduction – Infrared Dark Cloud (IRDC)

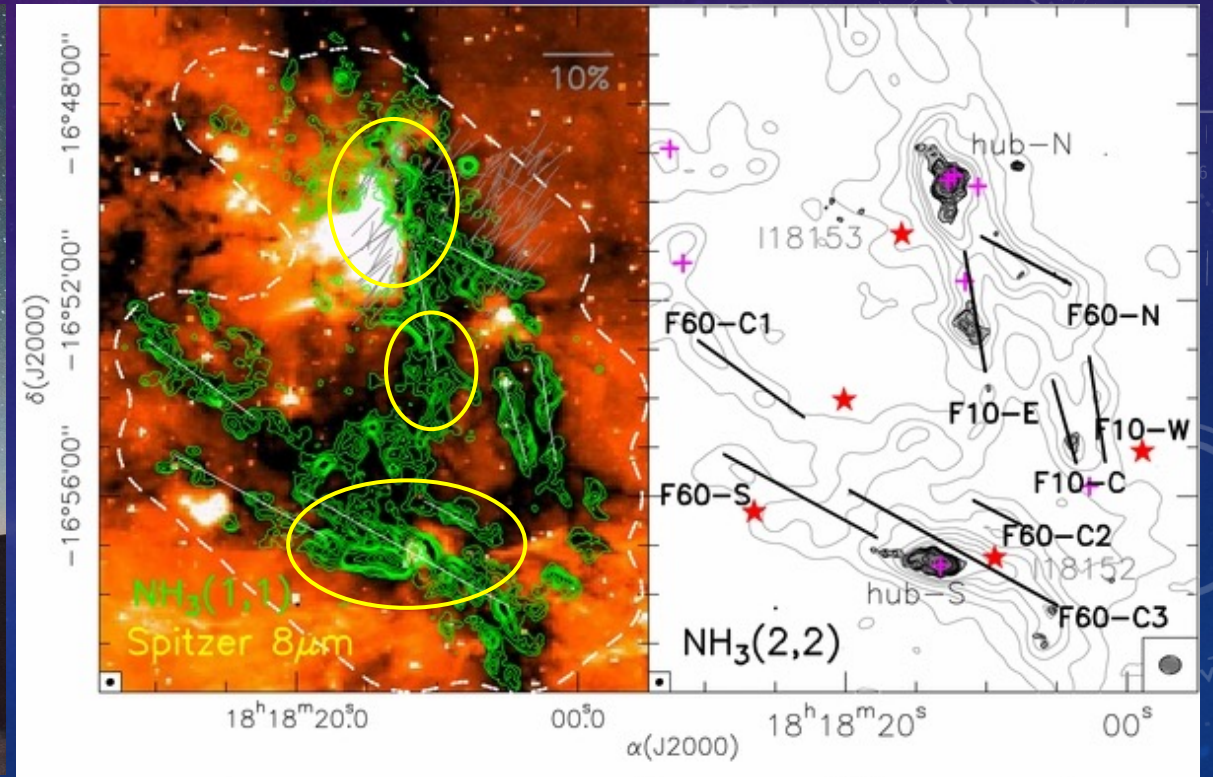
- Exhibit significant mid-IR opacity
- Extreme properties
 - Cold (< 20 K)
 - Enormous column densities ($> 10^{23} - 10^{25} \text{ cm}^{-2}$)
 - Some are dark at 7 to 100 μm
- Represent early stages of star formation



Perault et al. 1996; Egan et al. 1998; Carey et al. 1998, 2000; Hennebelle et al. 2001

Introduction - Observation data

- IRDC G14.225-0.506 by ALMA with mosaic and single pointing
- 1.3 mm dust continuum emission; spectral line emission (CO, N₂D⁺, etc.)



Credit: ESO/C. Malin (christophmalin.com)

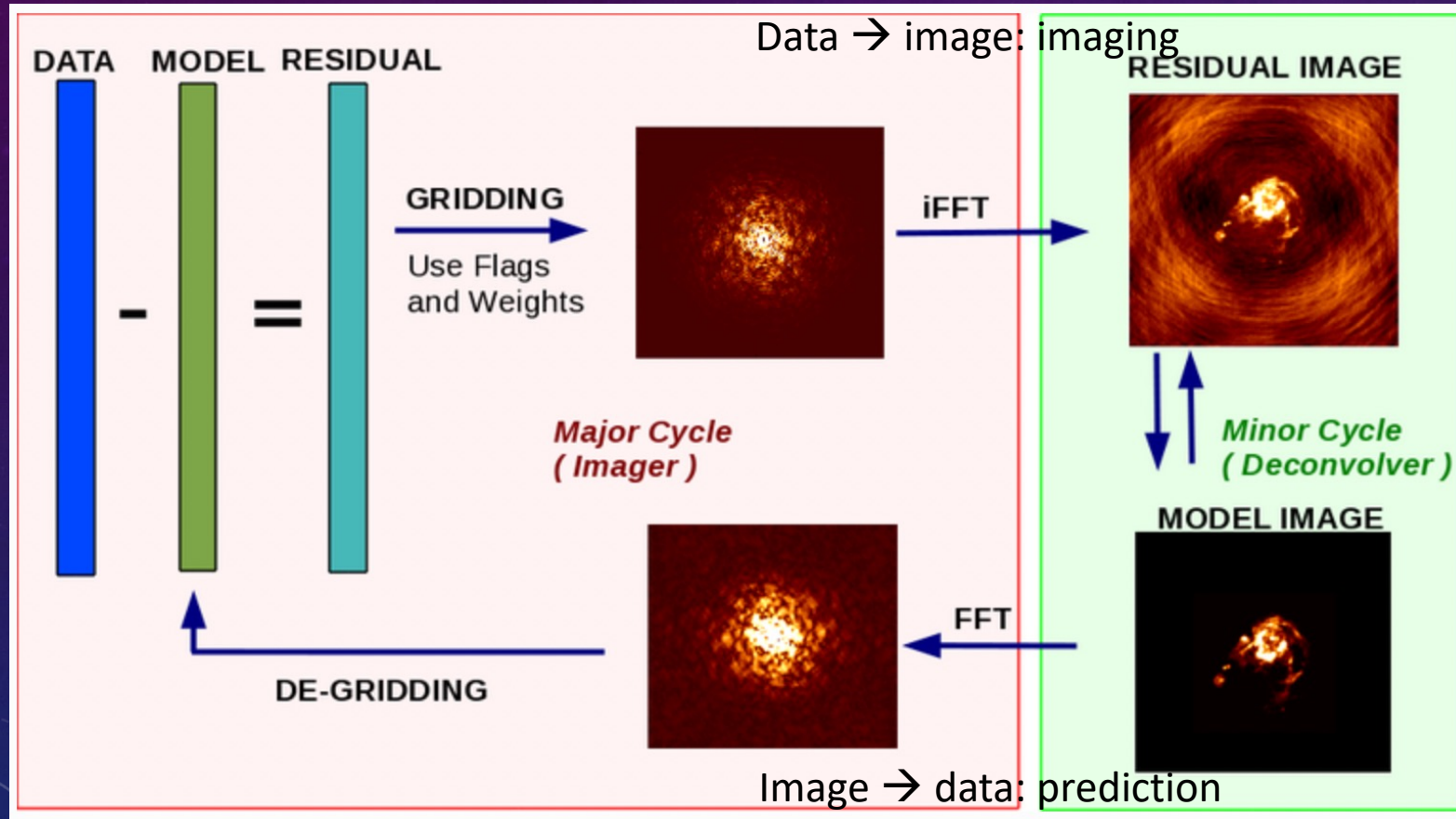
Busquet, Zhang et al. 2012

Method – Imaging by CASA

- Continuum subtraction → Cleaning → Images

Van Cittert-Zernike Theorem:

$$V_{\nu}(u, v) = \iint I_{\nu}(l, m) e^{-2i\pi(ul+vm)} dl dm$$

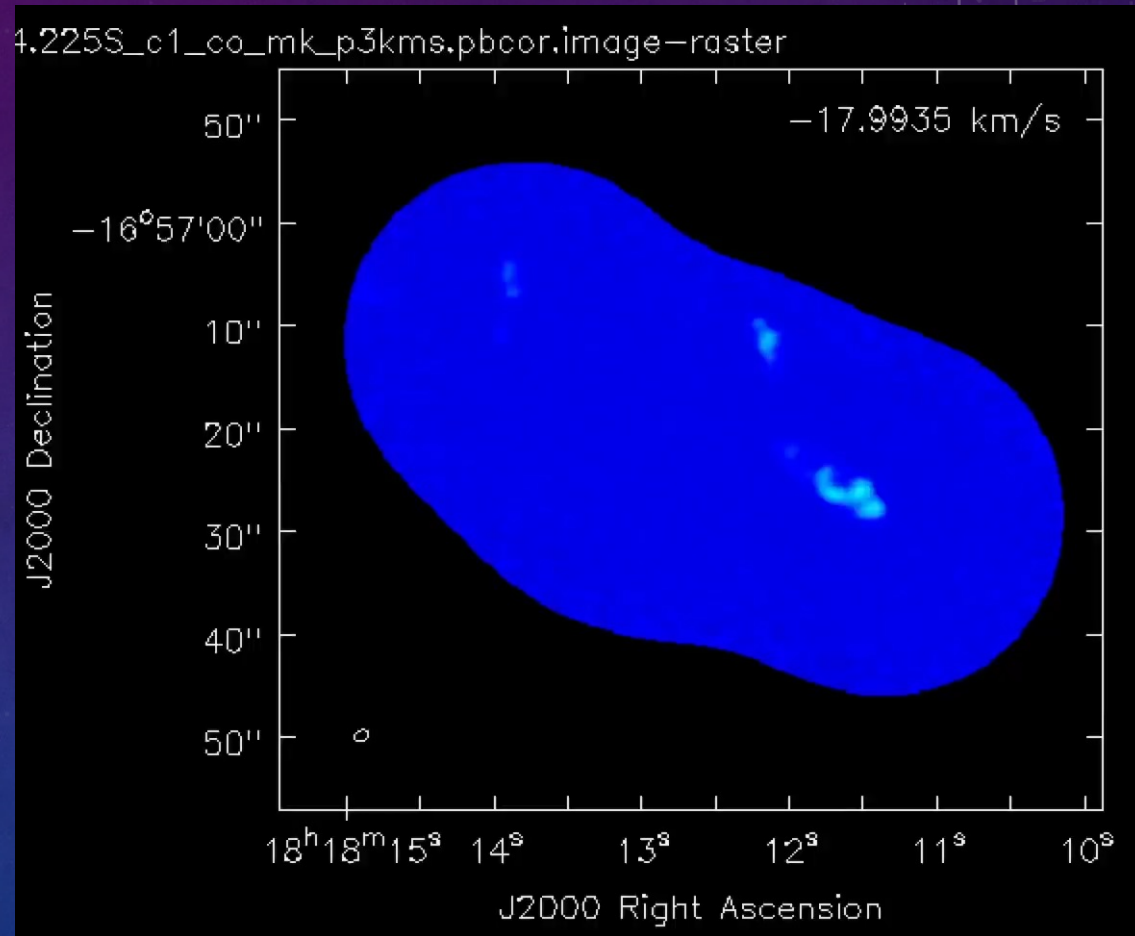


“true” brightness
↓
 $I_{\nu}^D(l, m) = I_{\nu}(l, m) \otimes B(l, m)$
↑ ↑
Dirty image Dirty Beam

Ay191 Fundamentals of interferometry, Qizhou Zhang & Eric Koch, 2022

Results - Images

- A taste of CO image:
 - intensity map from spectral line window
 - velocity channels ranging from around -18 km/s to 57 km/s
 - further analysis:
spatially confined wings and structures appeared in consecutive channels identified as outflows



What can we learn from the images?

- Outflows identified from line emissions (e.g., CO, SiO, H₂CO, etc.):
 - outflow parameters, e.g., mass, momentum, energy, dynamical timescale
 - as an implication for high mass star formation by accretion
- Dense cores studied from continuum data:
 - identified by dendrogram
 - mass functions analysis, e.g., power-law index

(Ohashi et al. (2016), Zhang et al. (2005))

Conclusion

- The infrared dark cloud is cold, dense, and dark.
- Study on the infrared dark clouds (IRDC) could reveal structures and properties of filaments, dense cores, outflows, etc.
 - better understand the early-stage formation of stars

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